Nuclear Fission in r-Process Nucleosynthesis

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Collaborators

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Introduction



Motivation

- We don't know where the r-process occurs.
 Supernova? Neutron star merger? Gamma-ray burst?
- If the fission occurs, we can expect that the elemental abundance pattern ejected from the site significantly changes.

Input data(1) Fission Fragment Mass Distribution



Our using data of the fluid

- neutrino-driven supernova(Woosley et al.1994) initial ye=0.43, initial density~10⁶[g/cm³]
- magnetohydrodynamic jet explosion of supernova(Nishimura et al.2006) initial ye=0.16, initial density~10⁷[g/cm³]
- neutron star merger (Freiburghaus et al.1999) initial ye=0.10, initial density~10¹¹[g/cm³]

topic

- How dose the fission affect the final abundance pattern for each r-process candidate site using new fission data?
- 2. How many fission cycles occur.

Neutrino Driven Wind



abundance

mass number

MHD Jet



abundance

Neutron Star Merger



Result

- The final abundance patterns of the neutrinodriven wind and the MHD jet are not affected by the nuclear fission, while the neutron star merger is largely affected by it.
- Therefore, we consider only the neutron star merger to investigate how many fission cycles occur.

topic

- How does the fission affect the final abundance pattern for each r-process candidate site using new fission data?
- How many fission cycles occur?

Indicator of fission



Indicator of fission



Result

- Almost all the nuclei fissioned.
- Fission cycle started but neutron capture stopped before one cycle is finished.

Summary

- We calculated r-process nucleosynthesis for three candidate site(Supernova ,MHD jet ,Neutron star merger) with new fission data.
- We could find that there is the effect of fission for Neutron star merger only.
- We formed the quantity to count fission cycles and there are no fission cycles even for neutron star merger.

Future Work

• We will apply our network code to modern numerical simulation of neutron star merger(Korobkin et al. 2012).